

30 **Abstract**

31 To constrain the Mg isotopic composition of the oceanic mantle, investigate Mg
32 isotope fractionation of abyssal peridotites during seafloor alteration, and assess Mg
33 budget in the oceans, a suite of 32 abyssal peridotite samples from the Gakkel Ridge
34 and Southwest Indian Ridge (SWIR) was, for the first time, selected for
35 high-precision Mg isotope analyses. Although most of these samples are extensively
36 altered, largely by serpentinization and weathering, primary olivine, diopside and
37 enstatite grains are preserved in some samples. Olivine grains from the least altered
38 samples have $\delta^{26}\text{Mg}$ varying from -0.30 to -0.12 ‰ ($n=7$), whereas enstatite and
39 diopside have $\delta^{26}\text{Mg}$ varying from -0.27 to -0.16 ‰ ($n=7$), and from -0.23 to -0.09 ‰
40 ($n=6$), respectively. Whole-rock $\delta^{26}\text{Mg}$ values range from -0.24 to 0.03 ‰ with an
41 average of -0.12 ± 0.13 ‰ (2SD, $n=32$). Strongly serpentinized peridotites have lower
42 average $\delta^{26}\text{Mg}$ values ($\delta^{26}\text{Mg} = -0.19 \pm 0.07$ ‰, 2SD, $n=7$) than
43 weathering-dominated ones ($\delta^{26}\text{Mg} = -0.10 \pm 0.12$ ‰, 2SD, $n=25$). Calculated Mg
44 isotopic compositions of fresh mantle peridotites vary from -0.29 to -0.13 ‰, beyond
45 the previously reported range of the subcontinental lithospheric mantle ($-0.25 \pm$
46 0.04 ‰) and the analytical uncertainty (± 0.07 ‰, 2SD). Our study therefore indicates
47 that the oceanic mantle may have similar but slightly heterogeneous Mg isotopic
48 compositions to that of subcontinental lithospheric mantle. Secondary serpentinization
49 does not fractionate Mg isotopes of abyssal peridotites, whereas low-T weathering
50 and formation of clay can result in the enrichment of heavy Mg isotopes in abyssal
51 peridotites. This study also demonstrates that fluid-rock interaction does not

52 necessarily produce rocks with intermediate Mg isotopic compositions. Magnesium
53 isotopes of the rocks thereafter are dependent on the secondary minerals formed. We
54 also conclude that the release of light Mg isotopes into the ocean during alteration of
55 abyssal peridotites can be an important influx of Mg for the seawater Mg budget.
56 Abyssal peridotites with a heavy Mg isotopic signature can be recycled into the
57 mantle in subduction zones and may thus result in heterogeneous Mg isotopic
58 compositions of the oceanic mantle and heavy Mg isotopic compositions of arc
59 magmas.

60

61 **Keywords:** abyssal peridotite, magnesium isotope, mantle heterogeneity, magnesium
62 cycling, seafloor alteration

63

64 **1. Introduction**

65 Abyssal peridotites comprise an important part of the oceanic crust and thus have
66 experienced prolonged fluid-rock interactions before being recycled as one of the
67 components into subduction zones (e.g., [Snow and Dick, 1995](#); [Mével, 2003](#); [Bach et al., 2004](#); [Li and Lee, 2006](#); [Paulick et al., 2006](#)). The fluid-peridotite interaction has
68 important consequences on the chemical budgets of oceans ([Thompson and Melson, 1970](#); [Snow and Dick, 1995](#)), and is believed to be dominated by hydration reactions
69 of olivine and pyroxenes that lead to formation of serpentine at a relatively broad
70 range of temperatures (25 to 400 °C, [Klein et al., 2009](#)), or weathering, represented by
71 the formation of clay minerals at relatively low temperatures (e.g., [Snow and Dick,](#)

74 1995). Because abyssal peridotites are largely serpentinized and/or weathered oceanic
75 mantle rocks, their compositions are considered as complex records of the
76 overprinting processes of partial melting, melt-rock reaction, and melt refertilization
77 in the asthenosphere and lithosphere, followed by hydrothermal alteration and
78 weathering as they are uplifted and exposed on the seafloor (e.g., Hellebrand et al.,
79 2001; Craddock et al., 2013). Therefore, abyssal peridotites provide a window for
80 probing high-T mantle processes in relation with low-T hydrothermal alteration.

81 Magnesium is the most abundant element besides Si in abyssal peridotites.
82 Previous studies have indicated that serpentinization is not an isochemical process
83 with mass transfer of Ca, Na, Fe and Mn and decrease of MgO/SiO₂ ratio (Malvoisin,
84 2015). Marine weathering, on the other hand, is widely accepted to result in Mg loss
85 (Snow and Dick, 1995). Previous isotopic studies have shown that the Earth's
86 subcontinental lithospheric mantle has $\delta^{26}\text{Mg}$ values of -0.25 ± 0.04 ‰ (Teng et al.,
87 2010a). However, oceanic peridotites of Purang (Tibet) have slightly higher $\delta^{26}\text{Mg}$
88 values of -0.20 ± 0.10 ‰ (Su et al., 2015), and altered oceanic crust has mantle-like
89 Mg isotopic compositions with $\delta^{26}\text{Mg}$ ranging much wider from -2.76 to +0.21 ‰
90 (Huang et al., 2015; Teng, 2017). Magnesium isotopic composition of oceanic mantle
91 peridotites and the behavior of Mg isotopes during fluid-peridotite interactions in the
92 exposure of abyssal peridotites is poorly known so far. Studies of Mg isotopes of
93 terrestrial ultramafic rocks indicate that Mg isotope fractionation strongly depends on
94 the processes the rocks have undergone. Serpentinization was suggested not to
95 fractionate Mg isotopes whereas carbonation can result in isotopically heavy talc and

96 isotopically light magnesite (Beinlich et al., 2014). In addition, metasomatism of
97 mantle rocks is widely used to explain rocks with significant inter-mineral
98 fractionation to light Mg isotopic compositions (Xiao et al., 2013; Beinlich et al.,
99 2014; Wang et al., 2014; Hu et al., 2016a), whereas the influence of seawater on Mg
100 isotopes of oceanic basalts and abyssal peridotites so far was suggested to be very
101 limited due to the low Mg concentrations of seawater ($[Mg] = 53 \text{ mmol/l}$ or $MgO =$
102 $0.21 \text{ wt.}\%$) (Huang et al., 2015). However, a preliminary study by Wimpenny et al.
103 (2012) indicated that weathering of abyssal peridotites is comparable to weathering of
104 continental crust. This controversy mainly results from the limited database for Mg
105 isotopic compositions of abyssal peridotites.

106 To understand the Mg isotopic systematics of the abyssal peridotites and its
107 fractionation during seawater-rock interactions, we selected samples from four sites of
108 the Gakkel Ridge and SWIR for whole-rock and mineral Mg isotopic analyses, and
109 then calculated Mg isotopes of fresh oceanic mantle peridotites according to the
110 mineral percentage and Mg isotopes of the constituting minerals. Our studies show
111 that oceanic mantle may have slightly heterogeneous Mg isotopes and that significant
112 amount of light Mg isotopes has been released to the ocean through
113 seawater-peridotite interactions during the exposure of abyssal peridotites.

114

115 **2. Geological background**

116 The peridotites used in this study come from the Prince Edward (46.54 °S,
117 33.79 °E) and the Shaka Fracture Zones (FZs) (53.38 °S, 9.33 °E) on the

118 ultraslow-spreading Southwest Indian Ridge (SWIR), and three areas (84.64 °N,
119 4.22 °E, 84.83 °N, 4.66 °E and 85.44 °N, 14.52°E) in the Sparsely Magmatic Zone of the
120 ultraslow-spreading Gakkel Ridge (Fig. 1). Detailed studies of these abyssal
121 peridotites and their respective locations have been presented in previous publications
122 (e.g., Hellebrand et al., 2001; Thiede et al., 2002; Dick et al., 2003; Michael et al.,
123 2003; Liu et al., 2008).

124 The SWIR separates the African and Antarctic plates, and spreads at a full rate of
125 14 mm yr⁻¹, making it an ultraslow spreading ridge (Dick et al., 2003). Sixteen abyssal
126 peridotites in this study are from two areas of the SWIR collected on cruise 107 Leg 5
127 of the RV Atlantis II (AII107), and Protea Expedition, Leg 5 (Pr) of the RV Melville
128 from the Shaka and Prince Edward FZs, respectively (Dick et al., 1984). Peridotites of
129 the SWIR range from lherzolites to harzburgites and exhibit major element
130 heterogeneity in the initial mantle, preserved because of low degrees of melting
131 beneath the SWIR (Seyler et al., 2003). The Gakkel Ridge in the high Arctic Ocean
132 extends east for ~1800 km from the Lena trough, north of Greenland, toward the
133 continental margin of the Laptev Sea (Thiede et al., 2002). Full spreading rate
134 decreases west to east from 14.6 to 6.3 mm yr⁻¹, making the entire ridge ultra-slow
135 spreading (Dick et al., 2003). There are three distinct regimes with different relative
136 abundances of rock types: a western volcanic zone, a central sparsely magmatic zone
137 that is nearly amagmatic and an eastern zone of widely spaced volcanoes (Michael et
138 al., 2003). Fourteen abyssal peridotites from the Gakkel Ridge in this study are from
139 the central sparsely magmatic zone collected during the AMORE Expedition of the

140 USCGC Healy (HLY) and RV Polarstern (PS) (Thiede et al., 2002; Michael et al.,
141 2003). These peridotites include both depleted harzburgites and more fertile
142 lherzolites, and are inferred to have undergone low-degree melting in the current ridge
143 event due to the missing or overall thin crust along the ridge (Dick et al., 2003;
144 Michael et al., 2003). The Gakkel Ridge peridotites come from amagmatic to weakly
145 magmatic spreading segments, while the samples from SWIR are from large off-set
146 transforms that likely inhibited mantle melting due to the transform edge effect (e.g.,
147 Dick et al., 2010). The two suites from the SWIR, on the other hand, represent low
148 degree melting adjacent to the Shaka FZ, and moderately high degree melting
149 adjacent to the Prince Edward FZ. Given their overall range in composition from
150 diopside-poor harzburgite to diopside-rich lherzolite and the relatively thin crust at
151 both ridges, it is likely that their mantle sources had undergone previous melting
152 events prior to mantle melting during the current ridge event (e.g., Zhou and Dick,
153 2013).

154 Radiogenic isotopic studies (Sr, Nd, Pb and Os isotopes) of abyssal peridotites
155 have demonstrated their heterogeneity in both large and small scales, with similar
156 ranges to MORBs but extend to more depleted composition (e.g., Meyzen et al., 2005;
157 Liu et al., 2008; Warren et al., 2009). Iron stable isotopes of abyssal peridotites ($\delta^{56}\text{Fe}$
158 = 0.010 ± 0.007 ‰) are indistinguishable from chondrites (Craddock et al., 2013).
159 Debret et al. (2016) also indicated that ocean-floor serpentinization, represented by
160 the formation of lizardite, does not fractionate Fe isotope appreciably. On the contrary,
161 O and Li stable isotopes display signatures of interaction of abyssal peridotites with

162 serpentinizing fluids, leading to low $\delta^{18}\text{O}$ values (e.g. 3.7 to 2.6 ‰ for abyssal
163 peridotites from the Mid-Atlantic Ridge) (Agrinier and Cannat, 1997) and low $\delta^6\text{Li}$
164 values (e.g. -2.9 to -14 ‰ for abyssal peridotites from SWIR) (Decitre et al., 2002).

165

166 **3. Petrography**

167 Most of the abyssal peridotites from the SWIR and Gakkel Ridge are 60 to 90%
168 altered to clay and serpentine, while some are quite fresh, with only a few percent
169 serpentine and light olivine weathering (e.g., PS59 235-4, Fig. 1). The majority of our
170 samples are protogranular or porphyroclastic, together with some mylonites that have
171 undergone intense ductile deformation. The fresh samples contain coarse-granular
172 enstatite and diopside that are typically 2-5 mm in size. Olivine in the fresh samples is
173 well preserved with a microfracture network of black to brown alteration veins that
174 crisscross the olivine. These veins are filled with serpentine, Fe-oxides and oxidized
175 brown clay such as chlorite. Spinel occurs in minor amount and is mostly small and
176 fresh.

177 In the altered abyssal peridotites, rock color often indicates the degree of
178 serpentinization and/or weathering: orange to brown color indicates moderate to high
179 weathering of silicate minerals to clay (e.g. AII107 from Shaka FZ, Fig. 1), while
180 green and black colors indicate fresher, less weathered highly serpentinized rocks (e.g.
181 Pr from Prince Edward FZ, Fig. 1). In the serpentinization-dominated samples,
182 alteration minerals include serpentine, hornblende, andradite, chlorite and Cr-Fe
183 oxides (Fig. 2a), whereas in weathering-dominated ones, alteration minerals are

184 similar but there are more chlorite and clusters of microcrystalline Mg-Si-Al-Cr-Fe
185 minerals that resemble iddingsite (Fig. 2b-d), a typical weathering product of
186 mafic-ultramafic rocks. About 62% of the Polar Stern peridotites are weathered to
187 different degrees and only 35% are serpentinized (Thiede et al., 2002), similar to the
188 Healy samples (Michael et al., 2003). The Gakkel Ridge and Shaka FZ peridotites in
189 this study are dominated by weathering, resulting in orangish-brown to tan coloring,
190 with subordinate serpentine alteration, while peridotites from the Prince Edward FZ
191 have primarily undergone serpentinization, consistent with their green to black colors
192 in hand samples.

193

194 **4. Analytical methods**

195 Major element concentrations and loss-on-ignition (LOI) of the peridotites were
196 determined at Department of Geosciences, National Taiwan University, Taiwan.
197 Major element oxides were measured on fused glass disks using a Rigaku® RIX 2000
198 X-ray fluorescence (XRF) spectrometer. LOI was obtained by routine procedures.
199 Analytical uncertainties were estimated to be better than 5% (relative) for XRF
200 analyses (Lee et al., 1997).

201 All chemical procedures including sample dissolution, column chemistry and
202 instrumental analysis of Mg isotopes were conducted at the Isotope Laboratory of the
203 University of Washington, Seattle, following the procedures of previous studies (Teng
204 et al., 2007, 2010a, 2015). Weathered/alteration rinds and carbonate veins were cut off
205 prior to crushing. Mineral separates were cleaned three times (3×10 min) in Milli-Q

206 water with an ultrasonic bath. Sample dissolution and column chemistry were carried
207 out in the clean lab. Mineral separates were cleaned in an ultrasonic bath to remove
208 alteration from grain boundaries and cracks. About 1 to 3 mg whole-rock powder or
209 mineral separates were dissolved in a mixture of Optima-grade concentrated
210 HF-HNO₃-HCl in sealed 7 ml Savillex Teflon[®] screw-top capsule on a 100 °C hot
211 plate in a laminar flow exhaust hood. After complete dissolution, the samples were
212 dried out and redissolved in 1 N HNO₃ before chromatographic separation. Chemical
213 separation of Mg was achieved by cation exchange chromatography using Bio-Rad
214 AG50W-X8 (200 to 400 mesh) resin in 1 N HNO₃. The same purification procedure
215 was performed twice to effectly remove matrix elements. The total procedural blank
216 was < 10 ng for Mg, representing <0.1% Mg loaded onto the columns. Detailed
217 processing methods can also be referred to [Hu et al. \(2016a\)](#) and references therein.

218 Magnesium isotopic compositions were analyzed using the sample-standard
219 bracketing method on a *Nu Plasma II* MC-ICPMS at the University of Washington
220 ([Teng and Yang, 2014](#)). Two standards, San Carlos olivine and Hawaiian seawater
221 which were processed together with the samples through column chemistry, were
222 analyzed with each batch of samples to monitor the accuracy and reproducibility. An
223 international standard of JB-1 basalt powder was also processed through the same
224 procedure with the rock samples to access the accuracy of Mg isotopic analyses.
225 Magnesium isotopic data are reported in delta (δ) notation in per mil relative to DSM3
226 standard (a solution in 3% HNO₃ made from pure Mg metal, [Galy et al., 2003](#)): δ²⁶Mg
227 (‰) = [(²⁶Mg/²⁴Mg)_{sample}/(²⁶Mg/²⁴Mg)_{DSM-3}-1] × 1000. Results for these standards are

228 in agreement with previously published data (Table 1). Repeated analyses indicate
229 data reproducibility is better than ± 0.07 ‰ (2 standard deviations, SD) for $\delta^{26}\text{Mg}$
230 (Teng et al., 2015).

231

232 5. Results

233 5.1 Whole-rock major element compositions

234 The peridotites in this study show large variations in SiO_2 (32.4 to 51.8 wt.%),
235 MgO (25.7 to 46.5 wt.%), Fe_2O_3 (6.80 to 12.6 wt.%), Al_2O_3 (0.17 to 5.22 wt.%), TiO_2
236 (0 to 0.17 wt.%), and CaO (0.49 to 16 wt.%), with LOI varying from 0.9 to 18.2 wt.%
237 (Appendix 1). Recalculated to volatile free compositions, the average compositions of
238 SiO_2 (44.0 ± 2.8 1σ wt.%), MgO (39.6 ± 3.9 wt.%), Fe_2O_3 (10.3 ± 1.4 wt.%), Al_2O_3
239 (1.9 ± 1.1 wt.%), TiO_2 (0.07 ± 0.04 wt.%) and CaO (4.0 ± 4.5 wt.%) range much
240 larger than for fresh or reconstructed abyssal peridotite compositions, which have
241 quite restricted ranges for most major elements (Snow and Dick, 1995). In particular,
242 our samples have on average lost Mg, gained Al and Ca.

243 5.2 Whole-rock Mg isotopic compositions

244 Magnesium isotopic compositions of reference samples and abyssal peridotites
245 are reported in Table 1. $\delta^{26}\text{Mg}$ values of abyssal peridotites display wide variations
246 from -0.24 to 0.03 ‰ with an average value of -0.12 ± 0.13 ‰ (2SD, $n = 32$) (Fig. 3).
247 In particular, rocks from the Prince Edward FZ (Pr 18 and Pr 19) that undergone
248 intense serpentinization with less weathering have, on average, lighter Mg isotopic
249 compositions ($\delta^{26}\text{Mg} = -0.24$ to -0.14 ‰ with an average of -0.19 ± 0.07 ‰, 2SD, $n =$

250 7) than abyssal peridotites from other localities ($\delta^{26}\text{Mg} = -0.21$ to 0.03 ‰ with an
251 average of -0.10 ± 0.12 ‰, 2SD, $n = 25$). These values are slightly heavier than
252 average values of oceanic basalts and mantle peridotites ($\delta^{26}\text{Mg} = -0.25 \pm 0.07$ ‰)
253 reported in previous studies (e.g., [Handler et al., 2009](#); [Yang et al., 2009](#); [Teng et al.,](#)
254 [2010a](#); [Huang et al., 2011](#); [Lai et al., 2015](#))

255 **5.3 Mineral magnesium isotopic composition**

256 Magnesium isotopic compositions of olivine, enstatite and diopside from seven
257 relatively fresh abyssal peridotites vary in a small but discernible range. Olivine and
258 enstatite have $\delta^{26}\text{Mg}$ values from -0.30 to -0.12 ‰ and -0.27 to -0.16 ‰, respectively,
259 slightly lower than that of diopside ($\delta^{26}\text{Mg} = -0.23$ to -0.09 ‰) ([Table 2](#)). Although
260 slight variation exists for the same mineral from different samples, the $\delta^{26}\text{Mg}$ values
261 of olivine and enstatite are similar and within the mantle Mg isotopic range, slightly
262 lower than that of diopside, consistent with existing published data ([Handler et al.,](#)
263 [2009](#); [Yang et al., 2009](#); [Huang et al., 2011](#); [Liu et al., 2011](#); [Xiao et al., 2013](#); [Hu et](#)
264 [al., 2016b](#)) and equilibrium theoretical predictions ([Schauble, 2011](#); [Huang et al.,](#)
265 [2013](#)). Whole-rock Mg isotopic compositions for the primary mantle peridotites are
266 calculated according to mineral percentage and $\delta^{26}\text{Mg}$ values of the constituting
267 minerals in order to eliminate the effect of late-stage alteration along the cracks of
268 olivine. The calculation results show that mantle peridotites have slightly varied
269 whole-rock $\delta^{26}\text{Mg}$ values from -0.29 to -0.13 ‰ ([Table 2](#)), beyond the analytical
270 uncertainty (± 0.07 ‰) acquired in our lab.

271

272 **6. Discussion**

273 The heavy Mg isotopic composition of abyssal peridotites compared to Earth's
274 mantle ($\delta^{26}\text{Mg} = -0.25 \pm 0.04 \text{ ‰}$, [Teng et al., 2010a](#)) is unexpected, considering its
275 reaction with isotopically light seawater ($\delta^{26}\text{Mg} = -0.83 \pm 0.09 \text{ ‰}$, [Ling et al., 2011](#)).
276 The diversity of samples with variable degrees of alteration allow us to discuss the
277 Mg isotopic composition of the oceanic mantle, the behavior of Mg isotopes during
278 seawater alteration, the implications of Mg isotope fractionation during
279 seawater-peridotite interaction for other fluid-rock interaction processes, and finally
280 the impact of this study on Mg isotopic recycling in the mantle and oceans.

281

282 **6.1 Whole-rock major element composition**

283 The large range in silicon and calcium content, recalculated to 100% volatile free,
284 indicates that both these elements are mobile during alteration. None of the minerals
285 in a generic peridotite have less than 40 wt.% SiO_2 , and the proportion of pyroxene to
286 olivine in abyssal peridotites is such that the SiO_2 content of fresh peridotite is
287 generally between 43 and 44 wt.%, while CaO is generally less than 3 wt.%. In our
288 samples, silica ranges from 32 to 52 wt.% and CaO from near zero to almost 20 wt.%
289 ([Fig. 4a,b](#)). Calcium versus silica also produces no significant correlation ([Fig. 4a](#)).
290 Aluminium has not remained constant, increasing to values much higher than seen in
291 any reconstructed abyssal peridotite composition, with up to 5.33 wt.% Al_2O_3 in
292 sample PS59 235-4. The average aluminium for the suite is at the top of the observed
293 range for reconstructed fertile lherzolite compositions.

294 From this, it is clear that alteration of our samples was not isochemical, and that
295 silica should not be assumed constant during serpentinization. Both aluminium and
296 titanium, however, are relatively immobile elements, and tend to be redistributed
297 among the mineral phases during alteration, rather than being lost or gained. This is
298 illustrated in [Figure 4c](#), where there is a good correlation between these elements.
299 Calcium and aluminium, with the exception of four outliers from the Gakkel Ridge
300 which were considered to have undergone carbonate alteration, also display a good
301 positive correlation ([Fig. 4b](#)). These correlations are consistent with residues of partial
302 melting, and the large range in degrees of melting inferred for these and other abyssal
303 peridotites from mineral modes and compositions (e.g., [Dick et al., 1984](#)). Based on
304 these correlations, the increase in aluminium in the altered peridotites likely reflects
305 its immobility rather than any actual addition, during alteration, hydration, and loss of
306 other species.

307 Four of the Protea Leg 5 samples have elevated TiO_2 and higher alumina than the
308 remaining three ([Fig. 4c](#)), all of which have relatively light $\delta^{26}\text{Mg}$ ([Fig. 4d](#)). The high
309 titanium can be explained in a number of ways, but one likely source is impregnation
310 by a MORB-like melt, which also means that plagioclase was present in the unaltered
311 peridotite ([Dick, 1989](#)). Generally, fresh plagioclase is not present in abyssal
312 peridotites, as it readily alters to prehnite and/or hydrogrossular \pm chlorite and clay.
313 Hence, this reaction could affect the Mg isotopic composition of the altered rock, and
314 could be a possible source of lighter isotopic compositions in some abyssal peridotites.
315 Otherwise, no consistent correlation can be seen between relatively immobile major

316 elements and the isotopic composition of our samples, consistent with there being no
317 relationship to igneous petrogenesis (e.g., [Fig. 4d](#)).

318

319 **6.2 Mantle Mg isotopic heterogeneity**

320 Previous studies on Mg isotopic composition of the Earth's mantle have analyzed
321 mineral separates and whole-rock compositions of mantle xenoliths from
322 subcontinental lithospheric mantle as representative of the Mg isotopic composition of
323 the upper mantle, or comparison to chondrites ([Handler et al., 2009](#); [Yang et al., 2009](#);
324 [Teng et al., 2010a](#); [Huang et al., 2011](#); [Liu et al., 2011](#); [Xiao et al., 2013](#); [Hu et al.,](#)
325 [2016a](#)). This indirect sampling is largely because of the inaccessibility of mantle rocks.
326 Although mantle xenoliths entrapped in lavas commonly display varied Mg isotopic
327 compositions due to metasomatism and/or contamination during their ascent to the
328 Earth's surface, particularly near subduction zones where intense fluid activities occur
329 (e.g., [Yaxley et al., 1991](#); [Ionov et al., 1993](#); [Rudnick et al., 1993](#)), relatively
330 homogeneous Mg isotopic compositions of -0.25 ± 0.04 ‰ for global peridotite
331 xenoliths from subcontinental lithospheric mantle have been reported by [Teng et al.](#)
332 [\(2010a\)](#). Fresh minerals of abyssal peridotites from oceanic mantle, however, are not
333 influenced by these processes. In addition, the fresh mineral separates are not affected
334 by serpentinization or weathering, and thus can provide Mg isotopic compositions of
335 the primary oceanic mantle. The seven relatively fresh abyssal peridotites from SWIR
336 and Gakkel Ridge in this study have homogeneous olivine, enstatite, diopside and
337 spinel compositions in each sample and within the range of typical mantle rocks

338 ([Appendix 2](#)). Mineral compositions of the coexisting silicates indicate they are fresh,
339 in equilibrium and have not been modified by late-stage processes.
340 Orthopyroxene-olivine and clinopyroxene-olivine Mg isotopic fractionations largely
341 fall along the theoretically calculated fractionation line by [Schauble \(2011\)](#) ([Fig. 5a](#),
342 [b](#)), indicating equilibrium Mg isotopic fractionation between the coexisting silicate
343 minerals for these mantle peridotites. Mantle olivine, however, has Mg isotopic
344 compositions varying from -0.12 to -0.30 ‰, beyond the analytical uncertainty (\pm
345 0.07 ‰). Similarly, whole-rock Mg isotopes of these mantle peridotites calculated
346 through mineral percentage and isotopic composition display heterogeneous variation
347 (-0.29 to -0.13 ‰) that is beyond the subcontinental mantle range ($\delta^{26}\text{Mg} = -0.25 \pm$
348 0.04 ‰, [Teng et al., 2010a](#)) and is not related to partial melting degrees represented by
349 Cr# of spinel ([Fig. 5c](#)). If Mg isotopes do not fractionate during partial melting and
350 melt extraction of the mantle rocks as previously demonstrated ([Handler et al., 2009](#);
351 [Yang et al., 2009](#); [Xiao et al., 2013](#); [Lai et al., 2015](#)) and this study shows, the Mg
352 isotopic signature of the olivine crystals and whole-rock mantle peridotites under the
353 Gakkel Ridge and the SWIR may represent heterogeneous oceanic mantle. This Mg
354 isotopic heterogeneity may result from the subduction and residence of oceanic
355 lithosphere in the mantle, which brings abundant MgO in isotopically heavy abyssal
356 peridotite and altered basalt and less MgO in isotopically light/heavy marine
357 sediments back into the convecting mantle. Similar phenomenon of O, Li and Fe
358 stable isotopic heterogeneity has already been reported in mantle rocks ([Krienitz et al.,](#)
359 [2012](#); [Williams and Bizimis, 2014](#)). Because of the large amount of Mg in abyssal

360 peridotites and altered basalts and the high density, their recycling into the mantle
361 may shift the Mg isotopes of mantle source to a heavier value and provide sources of
362 heavy Mg isotopes for arc magmas. This has just been identified by [Teng et al. \(2016\)](#)
363 in Lesser Antilles Arc lava, which displays a wide range of $\delta^{26}\text{Mg}$ (-0.25 to -0.10 ‰)
364 slightly heavier than the mantle value.

365

366 **6.3 Magnesium loss and Mg isotopic fractionation of abyssal peridotites** 367 **during seawater alteration**

368 Abyssal peridotites recovered from different localities of the mid-ocean ridges
369 represent exposed mantle rocks that have variably experienced post-magmatic
370 hydrothermal alteration (e.g., [Snow and Dick, 1995](#); [Humphris and Bach, 2004](#)).
371 Because of a relatively restricted $\delta^{26}\text{Mg}$ range of mantle source, any Mg isotopic
372 variation, if present, should reflect subsequent alteration process en route to and/or
373 exposed on the seafloor. The average $\delta^{26}\text{Mg}$ value of 32 abyssal peridotites is $-0.12 \pm$
374 0.13 ‰ (2SD, $n = 32$) and the Mg isotopic fractionation can be resolvable by up to
375 0.27 ‰, far beyond the analytical precision acquired (± 0.07 ‰). [Paulick et al. \(2006\)](#)
376 have summarized that an initial stage of high-temperature ($> 375 - 400$ °C)
377 hydrothermal interaction of peridotite with hydrothermal solutions results in
378 widespread serpentinization; the last stage of alteration is low-T (< 150 °C) interaction
379 of the serpentinites with ambient seawater that circulates in the near seafloor
380 environment, generating aragonite veinlets, Fe-oxyhydroxides and clays. Therefore,
381 serpentinization and seafloor weathering are evaluated separately to interpret Mg

382 isotope fractionation of abyssal peridotites.

383

384 **6.3.1 Serpentinization**

385 Serpentinites are abundantly exposed at slow and ultraslow spreading ridges,
386 largely produced by hydrothermal circulation into mantle rocks during faulting and
387 uplift to the seafloor and alteration of oceanic lithosphere (e.g., [Dick et al., 2003](#);
388 [Zhou and Dick, 2013](#); [Debret et al., 2016](#)). The fluid responsible for serpentinization
389 is seawater, possibly evolved by interaction with the crust. The extent of
390 serpentinization in the abyssal peridotites is commonly assessed from LOI values (e.g.,
391 [Craddock et al., 2013](#)). It is noteworthy that for a given LOI value, the MgO contents
392 vary in a large range of ~5 wt.% relative ([Fig. 3a](#)). Since marine weathering leading to
393 the formation of clay (hydrous aluminium phyllosilicates) can also introduce water
394 into the peridotites, the LOI value of abyssal peridotites cannot be simply considered
395 as an index of serpentinization. Petrographic observations should be carried out to
396 discriminate serpentinization and weathering before using simply geochemical
397 parameters.

398 Serpentine minerals, dominated by lizardite, chrysotile and antigorite, consist of
399 alternating infinite sheets of 4-coordinated Si and 6-coordinated Mg ([Mével, 2003](#); [Li
400 et al., 2014](#)). Among them, lizardite is a primary phase in seafloor serpentinites
401 because it forms at low temperature and pressure ([Mével, 2003](#)). It has been well
402 demonstrated that Mg isotopic fractionation is controlled by coordination and bonding
403 environment ([Liu et al., 2010, 2011](#); [Li et al., 2011, 2016](#); [Shauble, 2011](#); [Huang et al.,](#)

404 2013). Although the Mg isotopic composition of lizardite has not yet been reported, it
405 is speculated to be not significantly different from olivine, which also contains
406 6-coordinated Mg. This appears to be confirmed by our samples from Prince Edward
407 FZ (Pr in Fig. 3b), which have undergone intense serpentinization, have $\delta^{26}\text{Mg}$ values
408 within or minimally higher than the mantle range. Recent studies by Beinlich et al.
409 (2014) and Debret et al. (2016) also indicate that seafloor serpentinization does not
410 fractionate Mg and Fe isotopes, although increasing degrees of prograde
411 metamorphism resulting in the formation of antigorite may lead to loss of isotopically
412 light Fe to the fluids. Therefore, our study confirms that serpentinization is not likely
413 to fractionate Mg isotopes. We further propose that LOI cannot directly represent the
414 extent of serpentinization and that Mg isotopic fractionation in abyssal peridotites is
415 largely positively correlated with increasing volatile contents.

416

417 **6.3.2 Seafloor weathering**

418 Marine weathering of abyssal peridotites at the seafloor is known to affect
419 primary chemical compositions, leading to the loss of Mg from the rocks during this
420 process (Snow and Dick, 1995). Two types of low-temperature seafloor weathering
421 have been distinguished. One is dominated by weathering of olivine to iddingsite
422 (clay+goethite+aragonite), resulting in increase of Ca and loss of Si and Mg. The
423 other is characterized by weathering of serpentine and magnetite to talc + hematite ±
424 dolomite, with an increase in Si and Fe and a decrease in Mg (Humphris and Bach,
425 2004). The former has been identified petrographically as the primary weathering

426 process occurred in abyssal peridotites of this study (Figs. 1 and 2).

427 Seafloor weathering can be assessed by a plot of MgO/SiO₂ versus Al₂O₃/SiO₂ in
428 which weathering is characterized by departure from the terrestrial geochemical
429 fractionation array (Hart and Zindler, 1986). The proxy for weathering, MgO/SiO₂^{*},
430 is a quantitative measure of the departure of the measured peridotite MgO/SiO₂ ratio
431 from the primary ratio of each sample estimated from the MgO/SiO₂ versus
432 Al₂O₃/SiO₂ magmatic fractionation trend (Fig. 6a), assuming that Al and Si are
433 immobile during marine weathering (Snow and Dick, 1995). Our study shows that Mg
434 isotopes of variably weathered abyssal peridotites are slightly heavy with an average
435 δ²⁶Mg value of -0.12 ± 0.14 ‰ (2SD) compared with mantle peridotite value (δ²⁶Mg
436 = -0.25 ± 0.04 ‰, Teng et al., 2010a). As MgO/SiO₂^{*} ratios and the inferred degree of
437 weathering increase, the δ²⁶Mg ratios of abyssal peridotites increase as well (Fig. 6b).
438 This suggests that during seafloor weathering process, the Mg isotopic compositions
439 of the abyssal peridotites become heavy. Moreover, the negative correlation between
440 MgO content and δ²⁶Mg value indicates that heavy Mg isotope is enriched during
441 decrease of whole-rock MgO content (Fig. 6c). Therefore, the heavy Mg isotopic
442 compositions of abyssal peridotites could be due to the formation of iddingsite by
443 seafloor weathering. This is comparable to continental crust weathering, where clay
444 formation usually results in heavy Mg isotope enrichment compared with the mafic
445 protolith (e.g., Teng et al., 2010b; Huang et al., 2012; Liu et al., 2014).

446 The enrichment of heavy Mg isotope in residual clays and release of light Mg
447 isotope into the fluids during abyssal peridotite-fluids interactions can be modeled by

448 Rayleigh fractionation as it can for continental weathering. Using average MgO
449 concentration (44 wt.%, [Snow and Dick, 1995](#)) and $\delta^{26}\text{Mg}$ value in primary mantle
450 rocks, the apparent fractionation factor (α) obtained by modeling varies from ~ 1.0002
451 to 1.001 ([Fig. 7](#)), which corresponds to Mg isotopic fractionation between abyssal
452 peridotites and fluids from ~ 0.2 ‰ to 1 ‰. Accordingly, the fluids may contain Mg
453 isotopic compositions of -1.12 to -0.32 ‰. The large range in α value reflects a
454 complex process of dissolution of primary phases and precipitation of secondary
455 minerals and may therefore be a minimum ([Liu et al., 2014](#)).

456

457 **6.4 Implications for Mg isotopic fractionation during fluid-rock interactions**

458 The heavy Mg isotopic composition of abyssal peridotites relative to fresh
459 mantle peridotites indicates that during weathering of olivine, formation of clay
460 preferentially releases light Mg isotopes into seawater. This fractionation is mainly
461 driven by the difference in coordination environment between Mg in clay and
462 seawater. Mantle rocks with light Mg isotopes (lighter than mantle value) from
463 previous studies are usually explained as metasomatism by melts with light Mg
464 isotopic compositions ([Xiao et al., 2013](#); [Hu et al., 2016a](#)). By contrast, mantle rocks
465 with heavy Mg isotopes are commonly interpreted as due to kinetic isotope
466 fractionation by diffusion during melt-peridotite interaction ([Huang et al., 2011](#); [Xiao
467 et al., 2013](#)). Our study, however, indicates that rocks with heavier Mg isotopic
468 compositions do not need to react with melt/fluid that has high Mg content and heavy
469 Mg isotopic composition. Through reaction with isotopically light seawater, abyssal

470 peridotites can release light Mg isotopes into the ocean and retain heavy Mg isotopes
471 in the weathering products of clay. This indicates that the Mg isotopic composition of
472 a rock has no essential correlation with Mg content and Mg isotopic composition of
473 the reacting melt/fluid, but is strongly dependent on the secondary minerals formed.
474 Our study also indicates that rocks reacted with carbonate fluids with light Mg
475 isotopic signature do not necessarily become enriched in light Mg isotopes, which is
476 often explained as the reason for Mg-isotopically light mantle-derived rocks. Rather,
477 other evidence for reaction with carbonate fluids needs to be provided together and
478 the processes whether they were mixed or reacted and the secondary minerals formed
479 after reaction also need to be investigated (e.g. [Yang et al., 2012](#); [Huang et al., 2015](#);
480 [Wang et al., 2016](#)). If carbonate is formed as a secondary mineral resulting from
481 carbonate metasomatism, the rock is expected to have light Mg isotopic composition,
482 as has been shown by [Xiao et al. \(2013\)](#). Therefore, Mg isotopic composition of a
483 rock or mineral should be interpreted together with detailed petrographical
484 observation and identification of mineral assemblage.

485

486 **6.5 Implications for oceanic Mg isotopic budget**

487 Serpentinized peridotites are an important component of the oceanic crust at
488 slow and ultraslow spreading ridges. Yet riverine Mg is considered to represent the
489 main source for ocean's Mg budget ([Tipper et al., 2006](#)). Studies of Mg cycling in
490 oceans used to focus on the discrepancy of riverine light Mg isotopic input ($\delta^{26}\text{Mg} =$
491 -1.09% , [Tipper et al., 2006](#)) with global seawater of relatively heavier Mg isotopic

492 composition ($\delta^{26}\text{Mg} = -0.83 \text{ ‰}$, [Ling et al., 2011](#)). It is well recognized that
493 precipitation of dolomite preferentially incorporates light Mg isotopes, thus
494 maintaining the heavier Mg isotope budget of seawater compared to rivers ([De](#)
495 [Villiers et al., 2005](#); [Tipper et al., 2006](#)). In this study, we have shown that Mg
496 isotopes of abyssal peridotites fractionate during seafloor weathering and light Mg
497 isotope released by formation of isotopically heavy clays shifts the Mg isotope budget
498 of the ocean to an opposite direction. If we assume the primary mantle peridotite has
499 44% MgO ([Snow and Dick, 1995](#)) with $\delta^{26}\text{Mg} = -0.25 \text{ ‰}$ and the abyssal peridotite
500 contains 37% MgO with an average $\delta^{26}\text{Mg} = -0.12 \text{ ‰}$ (this study), the released Mg
501 (5%) has an Mg isotopic ratio of -1.31 ‰ . If the exposure of serpentinites could
502 represent $\sim 20\%$ of the seafloor as suggested by [Cannat et al. \(1995\)](#) and similar
503 weathering processes occurred on oceanic basalts and gabbros exposed on the seafloor,
504 the Mg released during weathering could shift Mg isotopes of the ocean significantly
505 to a lighter value. This indicates that deposition of light Mg isotopes from the ocean
506 has been significantly underestimated. Therefore, studies on global Mg cycling and
507 Mg isotopic budget of the oceans, in particular, need to take into account the release
508 of light Mg isotope into the ocean during seafloor weathering of abyssal peridotites.

509

510 **7. Conclusions**

511 Our study has, for the first time, reported high-precision Mg isotopic
512 compositions of abyssal peridotites from the Gakkel Ridge and SWIR. The main
513 conclusions to be drawn include:

514 1) Abyssal peridotites display variable $\delta^{26}\text{Mg}$ values ranging from -0.24 to
515 0.03 ‰, with an average of -0.12 ± 0.13 ‰ (2SD, $n = 32$), heavier than those of
516 subcontinental lithospheric mantle peridotites ($\delta^{26}\text{Mg} = -0.25 \pm 0.04$ ‰) reported by
517 [Teng et al. \(2010a\)](#). The heterogeneous and heavier Mg isotopic compositions reflect
518 Mg isotope fractionation during seafloor weathering, in which light Mg isotope was
519 preferentially released to the ocean whereas heavy Mg isotope was retained in clay
520 minerals, comparable to that of continental weathering on terrestrial planet.

521 2) The lack of Mg isotope fractionation in serpentinized abyssal peridotites
522 suggests that alteration by serpentinization probably does not fractionate Mg isotopes,
523 consistent with the study by [Beinlich et al. \(2014\)](#).

524 3) Fresh olivine grains and reconstructed fresh oceanic mantle peridotites have
525 $\delta^{26}\text{Mg}$ values from -0.30 to -0.12 ‰ and -0.29 to -0.13 ‰, respectively, beyond the
526 analytical uncertainty acquired (± 0.07 ‰) in our lab, indicating heterogeneous Mg
527 isotopic compositions of the oceanic mantle source.

528 4) Mg isotope fractionation during melt/fluid-rock interaction is largely
529 dependent on the secondary minerals formed, instead of $\delta^{26}\text{Mg}$ value of the
530 interacting melt/fluid. Therefore, detailed petrographical observation and
531 identification of mineral assemblage should be carried out before interpretation of Mg
532 isotopic signature.

533 5) Release of light Mg isotope into the seawater during weathering of abyssal
534 peridotite is an important influx of Mg to the ocean and influences its Mg content and
535 Mg isotopic budget, which may need to be compensated by simultaneous deposition.

536 Further studies should take this into account when constructing Mg isotopic budget
537 and evaluating Mg cycling of the oceans.

538

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551

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777 **Figure captions**

778 Fig. 1 Physiographic and geoid maps of (a) the Gakkel Ridge in Arctic Ocean and (b)
779 the SWIR in Indian and Atlantic Oceans with sample locations and petrographic
780 pictures showing olivine (Ol), enstatite (Opx), diopside (Cpx) and spinel (Sp) in
781 relatively fresh abyssal peridotites. Olivine sometimes can be altered along fractures.

782

783 Fig. 2 Back-scattered electron images showing various alteration minerals and

784 textures of the studied abyssal peridotites. (A) Primary minerals completely altered to
785 hornblende (Hbl), serpentine (Serp), chlorite (Chl) and andradite (And). (B) and (C)
786 Olivine (Ol) partially altered to serpentine and microcrystalline Fe-oxides and clay
787 (Fe-oxd+clay) aggregates. (D) Primary minerals altered to serpentine, hornblende and
788 chlorite.

789

790 Fig. 3 a-Bulk sample abyssal peridotite LOI values versus MgO contents showing a
791 negative correlation. b-Bulk sample abyssal peridotite $\delta^{26}\text{Mg}$ versus LOI values
792 showing a positive correlation for samples HLY and PS from Gakkel Ridge and 107
793 (represent AII107) from SWIR. Samples Pr from SWIR is largely within the Mg
794 isotopic range of mantle peridotites ($\delta^{26}\text{Mg} = -0.25 \pm 0.04\%$) given by [Teng et al.](#)
795 [\(2010a\)](#) shown in the grey band.

796

797 Fig. 4 Plots of (a) SiO_2 versus CaO , (b) CaO versus Al_2O_3 , (c) TiO_2 versus Al_2O_3 and
798 (d) TiO_2 versus $\delta^{26}\text{Mg}$ of abyssal peridotites recalculated to volatile free composition.
799 Symbols can be referred to [Fig. 3](#).

800

801 Fig. 5 Comparison of measured (a) orthopyroxene (Opx) - olivine (Ol) fractionation
802 and (b) clinopyroxene (Cpx) - olivine (Ol) fractionation of Mg isotopes with
803 theoretical calculation as a function of temperature, and (c) plot of Cr# of spinel
804 versus whole-rock Mg isotopes for the studied relatively fresh abyssal peridotites. The
805 grey band represents Mg isotopic range of subcontinental mantle peridotites given by
806 [Teng et al. \(2010a\)](#). Uncertainties of $\Delta^{26}\text{Mg}$ values quote 2SD and are calculated as
807 the square root of the sum of the square of individual errors for $\delta^{26}\text{Mg}$ values.
808 Theoretical calculation of Mg isotopic fractionation is according to [Schauble \(2011\)](#).

809

810 Fig. 6 a-Plot of $\text{Al}_2\text{O}_3/\text{SiO}_2$ versus MgO/SiO_2 used to quantitatively measure the
811 departure of the peridotite MgO/SiO_2 ratio from the primary ratio of each sample
812 estimated from the MgO/SiO_2 versus $\text{Al}_2\text{O}_3/\text{SiO}_2$ terrestrial geochemical fractionation

813 array (Hart and Zindler, 1986; Jagoutz et al., 1979), assuming that Al and Si are
814 immobile during marine weathering. b-Plot of $\delta^{26}\text{Mg}$ versus $\text{MgO}/\text{SiO}_2^*$ for abyssal
815 peridotites. $\text{MgO}/\text{SiO}_2^*$, the proxy for weathering, is a quantitative measure of the
816 departure of the measured peridotite MgO/SiO_2 ratio from the primary ratio of each
817 sample estimated from Fig. 6a. c-Plot of $\delta^{26}\text{Mg}$ versus MgO of abyssal peridotites
818 showing a broadly negative correlation. Grey area represents mantle peridotite Mg
819 isotopic compositions given by Teng et al. (2010a).

820

821

822 Fig. 7 Mg concentration normalized to primary Mg content ($\text{MgO} = 44 \text{ wt.}\%$) of fresh
823 mantle peridotite reconstructed by Snow and Dick (1995) versus $\delta^{26}\text{Mg}$ for abyssal
824 peridotites from Gakkel Ridge and SWIR. Curved lines illustrate Mg removal via
825 Rayleigh distillation for different values of the fractionation factor α ($\alpha =$
826 $(^{26}\text{Mg}/^{24}\text{Mg})_{\text{abyssal peridotite}} / (^{26}\text{Mg}/^{24}\text{Mg})_{\text{fluid}}$). Rayleigh distillation equation: $\delta^{26}\text{Mg}_{\text{abyssal}}$
827 $\text{peridotite} = (\delta^{26}\text{Mg}_{\text{mantle peridotite}} + 1000)f^{(1/\alpha-1)} - 1000$; f - the fraction of Mg remaining in
828 the rock, calculated from $\text{MgO}_{\text{mantle peridotite}} / 44$ (primary MgO content of mantle
829 peridotite); $\delta^{26}\text{Mg}_{\text{mantle peridotite}} = -0.25 \text{ ‰}$ (Teng et al., 2010a). Red square represents
830 composition of average mantle peridotites. The abyssal peridotites in this study are
831 plotted between $\alpha = 1.0002$ to 1.001 . Symbols can be referred to Fig. 3.

832

833

834 **Tables**

835 Table 1 Magnesium isotopic compositions of abyssal peridotites and standards

836

837 Table 2 Magnesium isotopic compositions of mineral separates and equilibrium
838 temperatures for the studied abyssal peridotites

839

840 **Appendix**

841 Appendix 1 Major element compositions of abyssal peridotites

842

843 Appendix 2 EPM mineral compositions of olivine, enstatite, diopside and spinel in the

844 studied abyssal peridotites

Table 1

Table 1 Magnesium isotopic compositions of abyssal peridotites and standards

Sample/standard	$^{26}\text{Mg}(\text{‰})$	2SD	$^{25}\text{Mg}(\text{‰})$	2SD	Sample/standard	$^{26}\text{Mg}(\text{‰})$	2SD	$^{25}\text{Mg}(\text{‰})$	2SD
<i>Gakkel Ridge (84.83°N, 4.66°E)</i>					<i>SWIR, Prince Edward FZ (46.54°S, 33.79°E)</i>				
HLY 4-37	-0.04	0.06	-0.04	0.04	Pr 18-12	-0.14	0.07	-0.06	0.05
HLY 4-37 rc	-0.05	0.06	-0.04	0.05	Pr 18-30	-0.18	0.07	-0.10	0.04
HLY 4-63	-0.14	0.07	-0.08	0.04	Pr 18-40	-0.21	0.07	-0.11	0.04
HLY 4-3	-0.01	0.07	0.01	0.04	Pr 19-4	-0.18	0.07	-0.10	0.04
HLY 4-29	-0.05	0.06	-0.01	0.04	Pr 19-5	-0.16	0.07	-0.09	0.04
<i>Gakkel Ridge (85.44°N, 14.52°E)</i>					Pr 19-47	-0.22	0.07	-0.11	0.04
HLY 40-26	-0.04	0.06	-0.02	0.05	Pr 19-96	-0.24	0.07	-0.11	0.04
HLY 40-53	-0.15	0.07	-0.08	0.04	Pr 19-96 ra	-0.21	0.05	-0.09	0.03
HLY 40-56	-0.18	0.06	-0.08	0.04	<i>Standards</i>				
HLY 40-43	-0.08	0.06	-0.04	0.04	San Carlos olivine	-0.24	0.06	-0.12	0.04
<i>Gakkel Ridge (84.64°N, 4.22°E)</i>					n=6	-0.22	0.07	-0.13	0.05
PS59 235-1	-0.14	0.06	-0.08	0.04		-0.23	0.07	-0.11	0.04
PS59 235-4	-0.14	0.07	-0.07	0.04		-0.22	0.08	-0.10	0.07
PS59 235-18	-0.13	0.06	-0.05	0.04		-0.21	0.07	-0.06	0.04
PS59 235-40	-0.10	0.07	-0.05	0.04		-0.19	0.06	-0.09	0.05
PS59 235-56	-0.09	0.06	-0.02	0.04	<i>Recommended</i>	-0.24	0.03	-0.12	0.02
PS59 235-82	-0.08	0.06	-0.03	0.04	Hawaiian seawater	-0.84	0.09	-0.44	0.06
PS59 235-67	-0.12	0.06	-0.06	0.04	n=5	-0.81	0.07	-0.43	0.05
PS59 235-68	-0.07	0.06	-0.03	0.04		-0.89	0.07	-0.44	0.04
<i>SWIR, Shaka FZ (53.38°S, 9.33°E)</i>						-0.79	0.08	-0.38	0.05
AII107 60-3	-0.07	0.06	-0.04	0.04		-0.88	0.06	-0.47	0.05
AII107 60-4	-0.15	0.07	-0.07	0.04		-0.83	0.09	-0.43	0.06
AII107 60-6	-0.09	0.07	-0.05	0.04	<i>Recommended</i>	-0.24	0.09	-0.12	0.06
AII107 60-68	-0.16	0.06	-0.07	0.04	JB-1 basalt	-0.24	0.09	-0.12	0.06
AII107 60-31	0.03	0.07	0.02	0.04	<i>Recommended</i>	-0.276	####	-0.146	###
AII107 60-31 ra	-0.01	0.06	-0.01	0.05	rc = repeat column chemistry from another				
AII107 60-31 rc	0.03	0.07	0.02	0.04	aliquot of dissolved sample solution;				
AII107 61-12	-0.03	0.07	-0.01	0.05	ra = repeat instrumental analysis on the same				
AII107 61-12 rc	-0.01	0.06	0.00	0.05	Mg-cut solution.				
AII107 61-82	-0.14	0.06	-0.08	0.04	FZ = Fracture Zone				
AII107 61-82 rc	-0.19	0.07	-0.11	0.04	Recommended Mg isotopic compositions of San				
AII107 61-98	-0.21	0.06	-0.11	0.04	Carlos olivine, Hawaiian seawater and JB-1				
AII107 61-98 ra	-0.20	0.05	-0.08	0.03	basalt are from Hu et al. (2016b), Ling et al.				
AII107 63-28	-0.08	0.06	-0.05	0.04	(2011) and Teng et al. (2015), respectively.				

Table 2

Table 2 Magnesium isotopic compositions of mineral separates and equilibrium temperatures for the studied abyssal peridotites

Sample	mineral	vol% ^a	$\delta^{26}\text{Mg}(\text{‰})$	2SD	$\delta^{25}\text{Mg}(\text{‰})$	2SD	Mg#	Cr#(Sp)	T (°C) ^b	$\delta^{26}\text{Mg}(\text{wr})^c$
Lherzolite										
PS59 235-1	Ol	84.4	-0.23	0.07	-0.09	0.04	90.8	0.14	1004	-0.22
	Opx	6.9	-0.17	0.07	-0.09	0.04	90.5			
	Cpx	8.6	-0.09	0.07	-0.05	0.04	91.0			
	Cpx ra	8.6	-0.06	0.05	-0.05	0.03				
PS59 235-4	Ol	59.0	-0.21	0.07	-0.11	0.04	89.7	0.3	1032	-0.22
	Opx	18.5	-0.27	0.07	-0.11	0.04	90.0			
	Cpx	22.5	-0.19	0.07	-0.12	0.04	89.9			
PS59 235-56	Ol	74.8	-0.18	0.07	-0.08	0.04	90.2	0.13	1065	-0.18
	Opx	20.7	-0.16	0.07	-0.10	0.04	90.1			
	Opx ra	20.7	-0.13	0.05	-0.09	0.03				
	Cpx	4.5	-0.16	0.07	-0.09	0.04	90.6			
PS59 235-82	Ol	56.5	-0.12	0.06	-0.05	0.05	91.1	0.21	1005	-0.13
	Ol ra	56.5	-0.10	0.05	-0.06	0.03				
	Opx	35.7	-0.17	0.06	-0.07	0.05	91.0			
	Cpx	7.8	-0.11	0.05	-0.06	0.03	92.1			
	Cpx ra	7.8	-0.12	0.06	-0.04	0.05				
HLY 4-63	Ol	70.3	-0.22	0.06	-0.12	0.05	89.7	0.38	1016	-0.22
	Opx	17.4	-0.17	0.06	-0.09	0.05	90.0			
	Cpx	12.3	-0.23	0.06	-0.12	0.05	90.7			
Cpx-bearing Harzburgite										
Pr 19-5	Ol	86.6	-0.30	0.06	-0.16	0.05	90.6	0.3	903	-0.29
	Opx	11.0	-0.20	0.06	-0.09	0.05	90.8			
	Cpx	2.4	-0.17	0.06	-0.09	0.05	91.8			
Harzburgite										
HLY 40-56	Ol	89.5	-0.19	0.06	-0.09	0.05	91.8	0.54	1041	-0.19
	Opx	10.5	-0.17	0.05	-0.08	0.03	92.1			

ra = repeat instrumental analysis on the same Mg-cut solution.

^a Volume estimates of each sample is calculated by point counting of $\sim 4 \times 2.5 \text{ cm}^2$ thin section.

^b T estimates of lherzolite and Cpx-bearing harzburgite are calculated according to Wells (1977), and those of harzburgite are according to Lindsley and Dixon (1976).

^c wr-whole rock; $\delta^{26}\text{Mg}(\text{wr})$ is calculated according to the volume, MgO content and Mg isotopic value of each constituting mineral.

Figure 1
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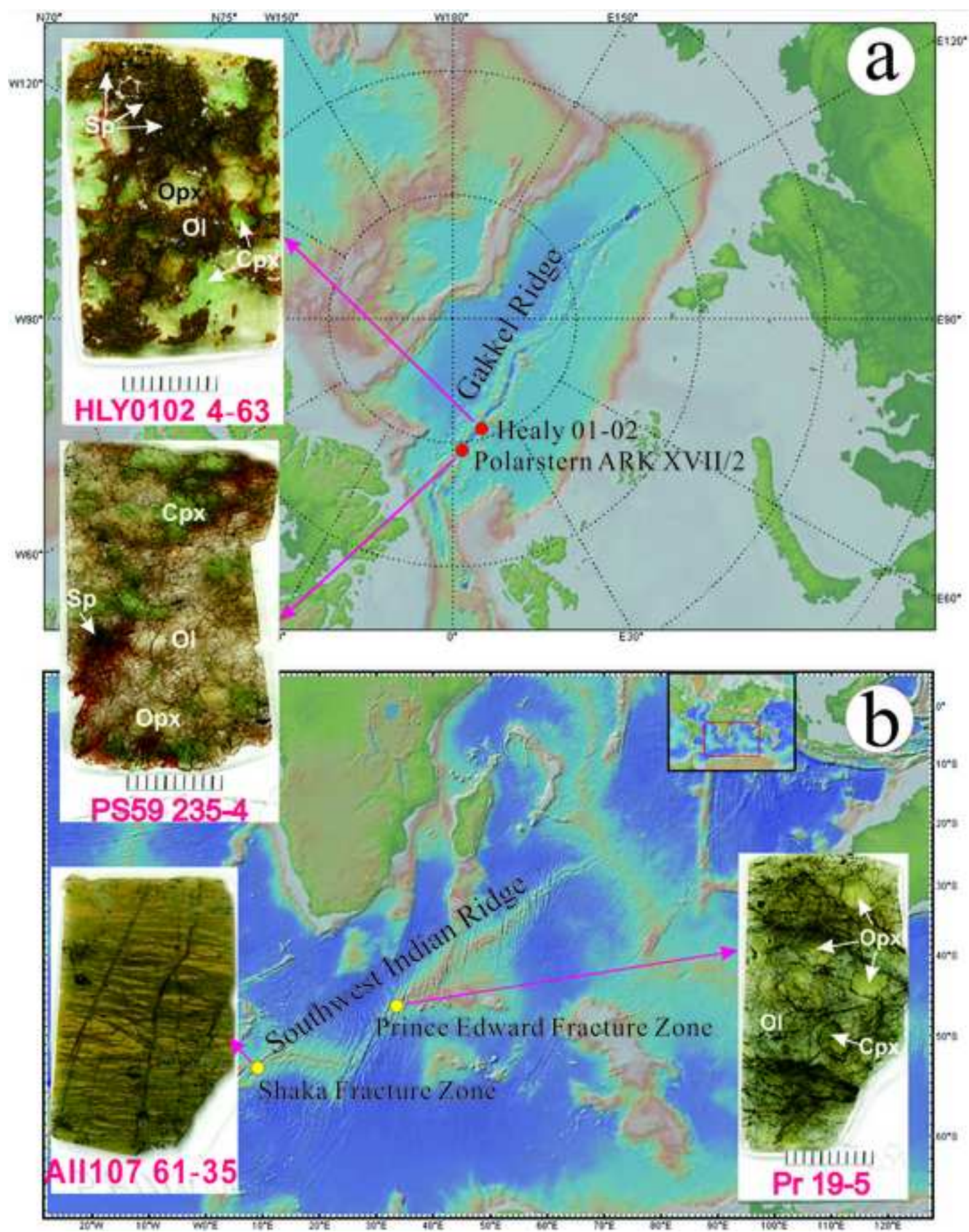


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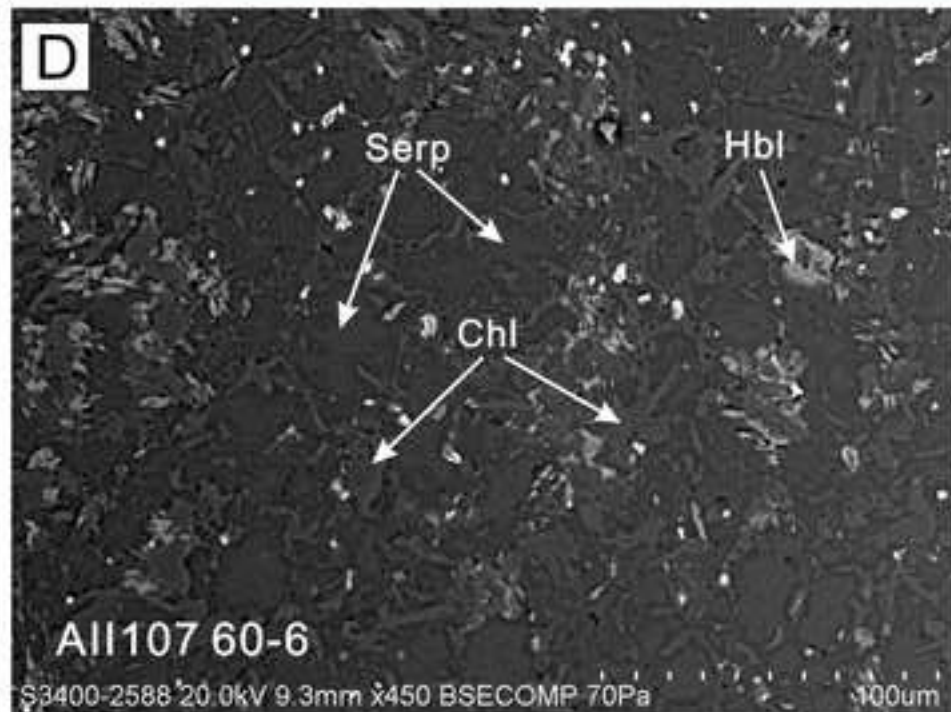
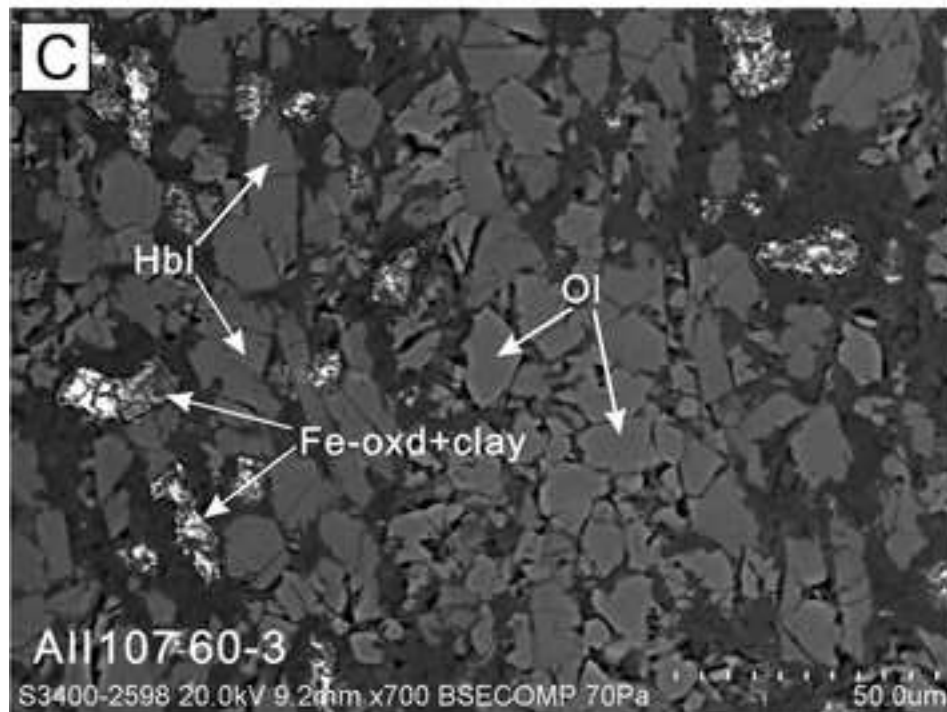
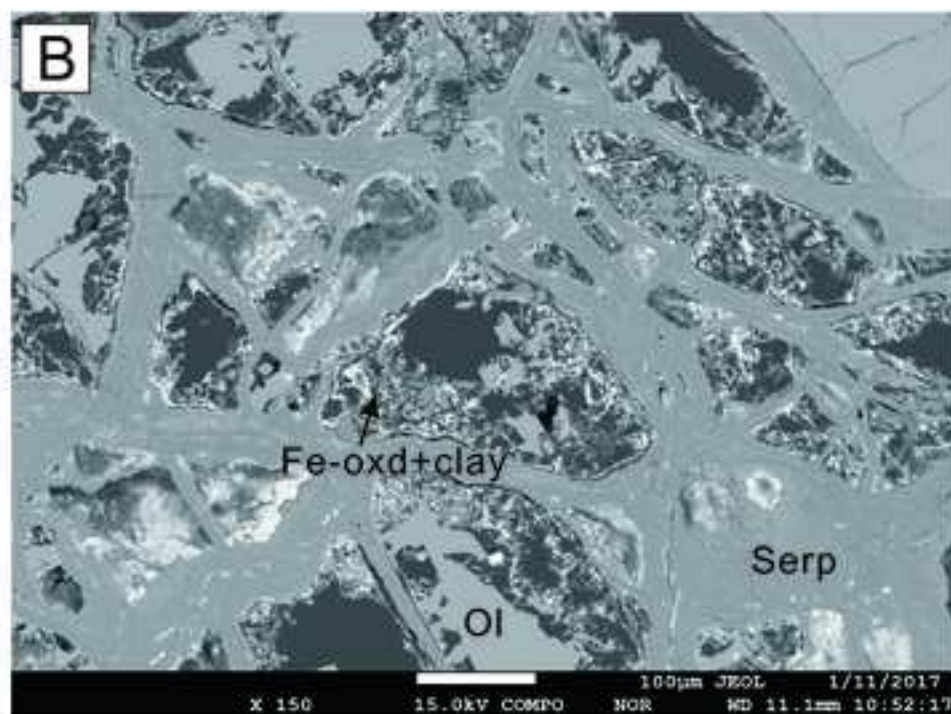
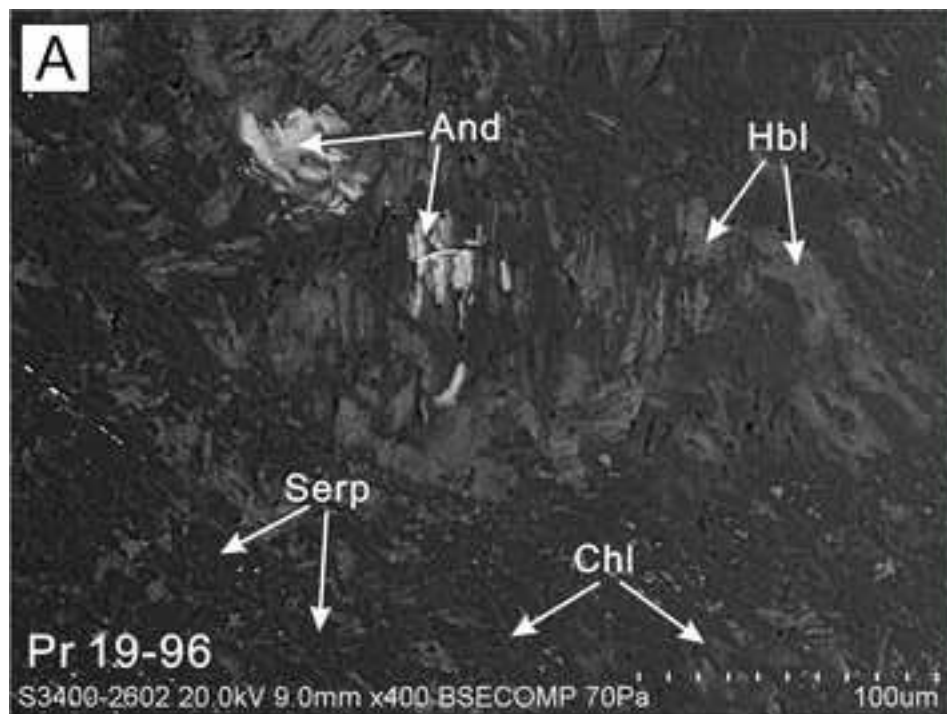


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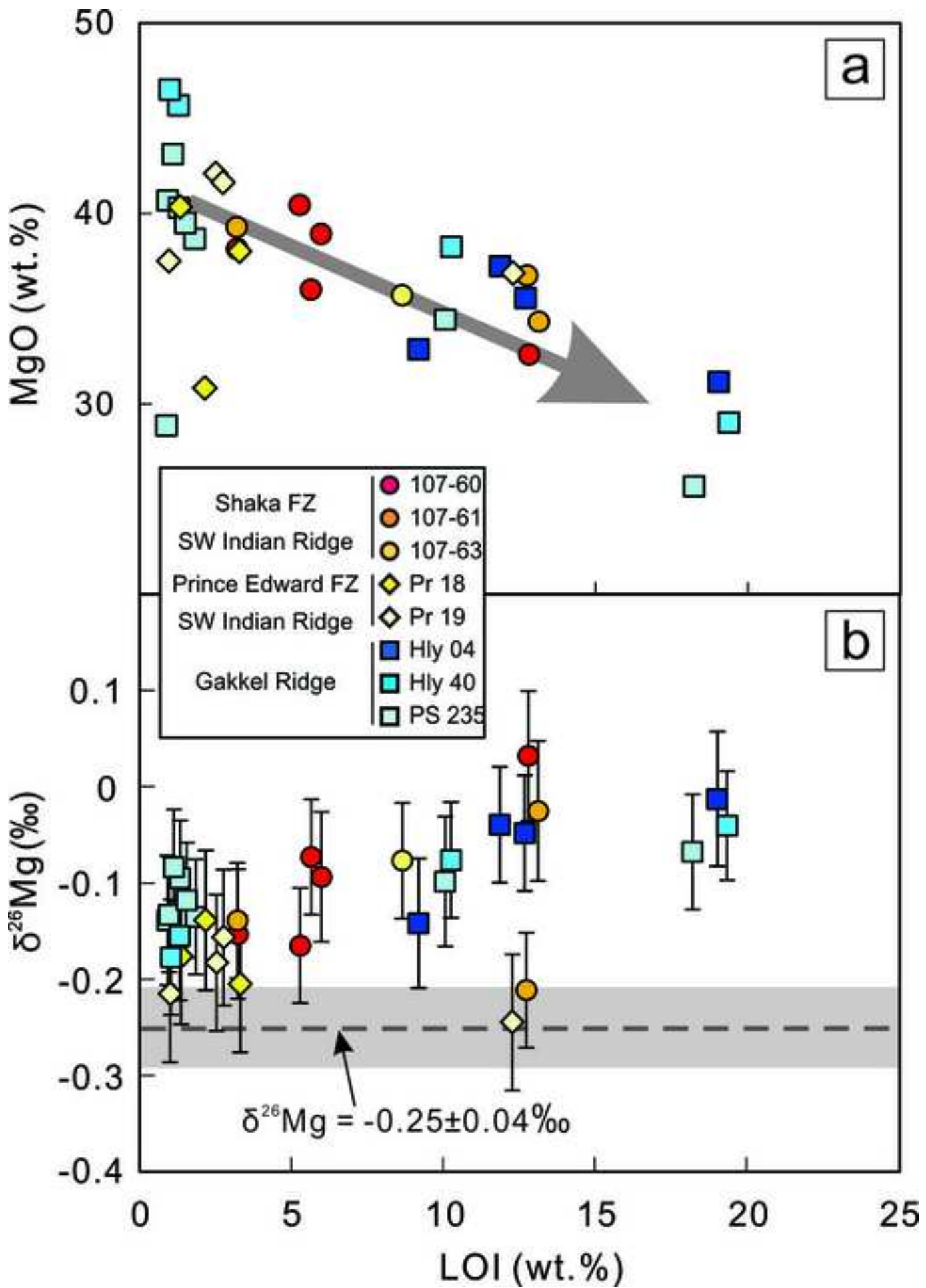


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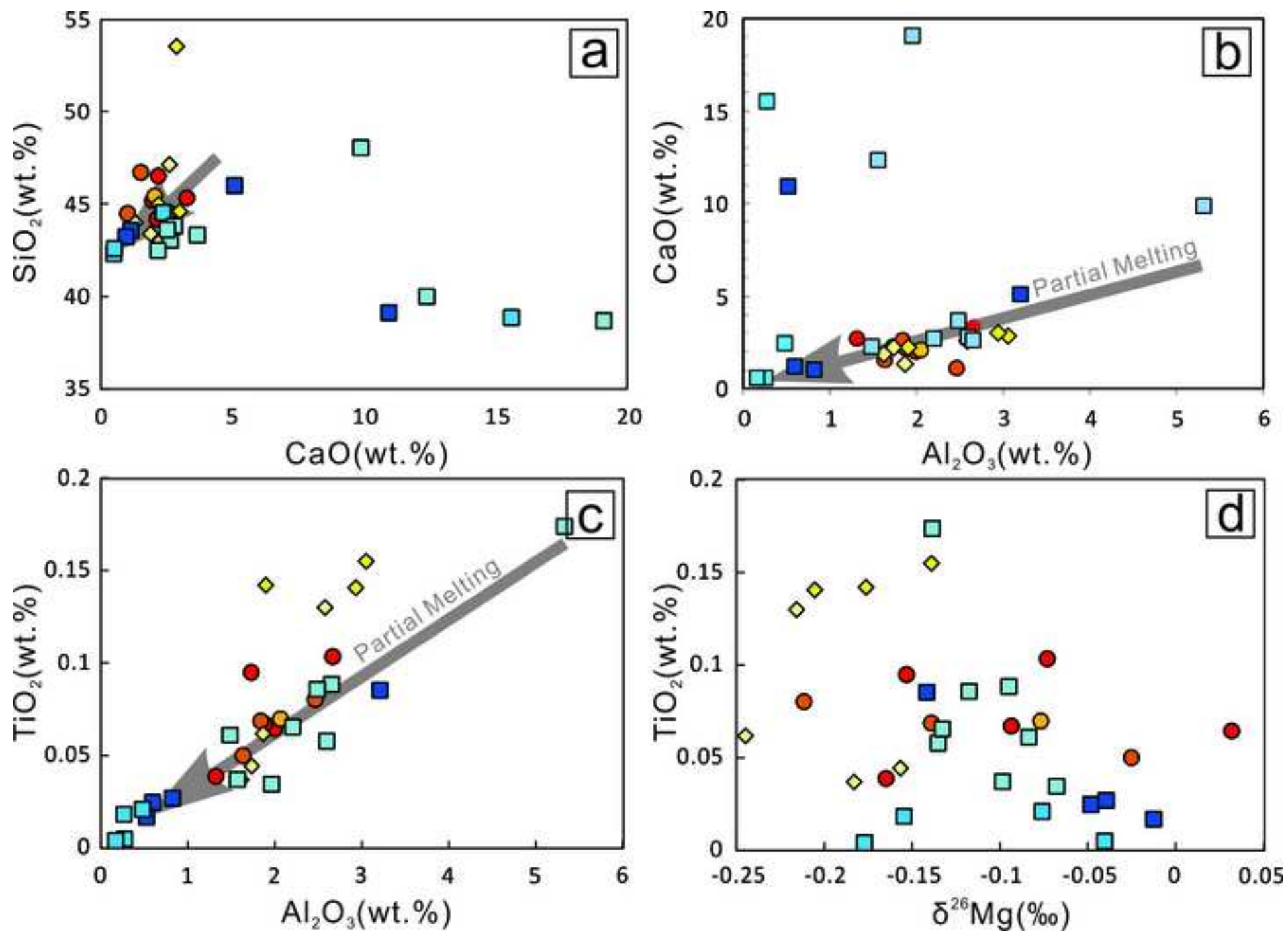


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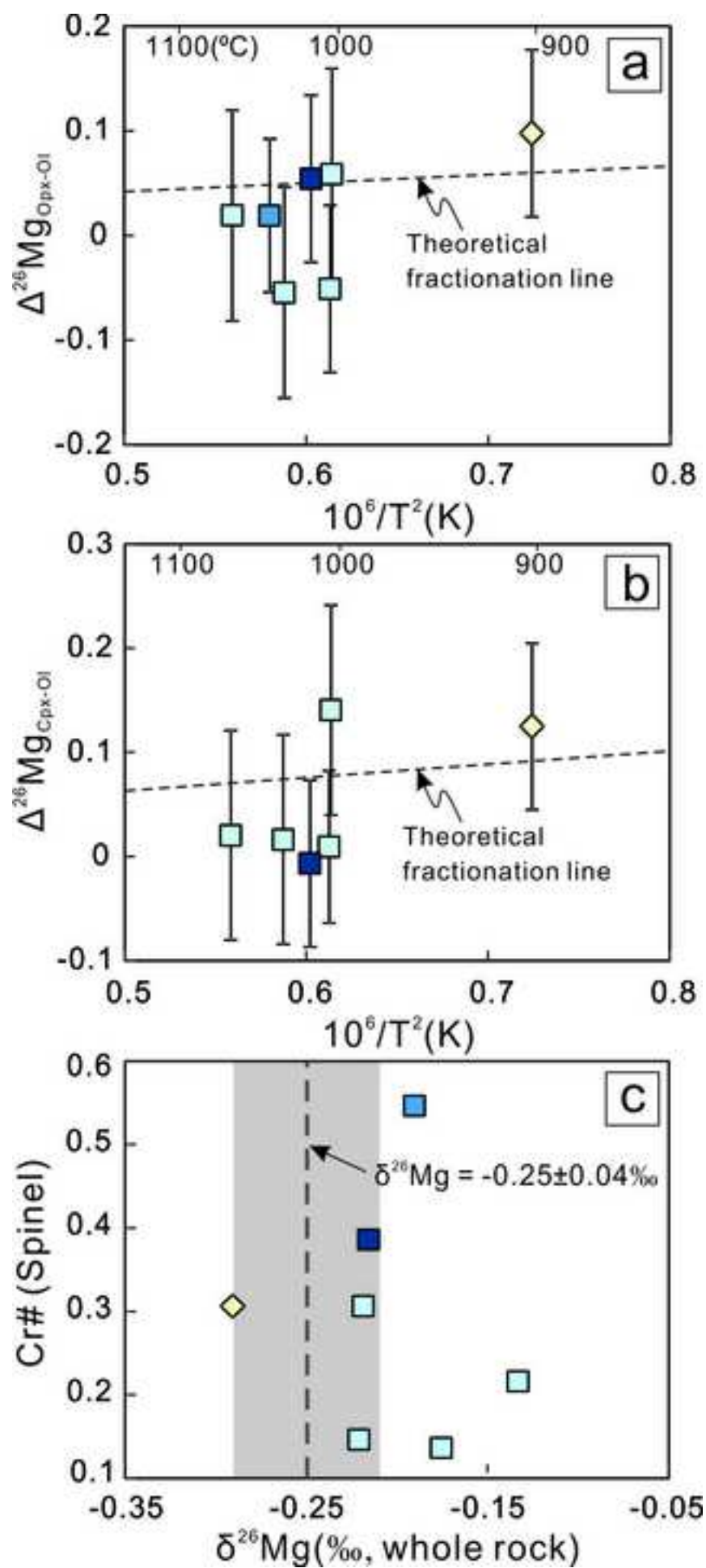


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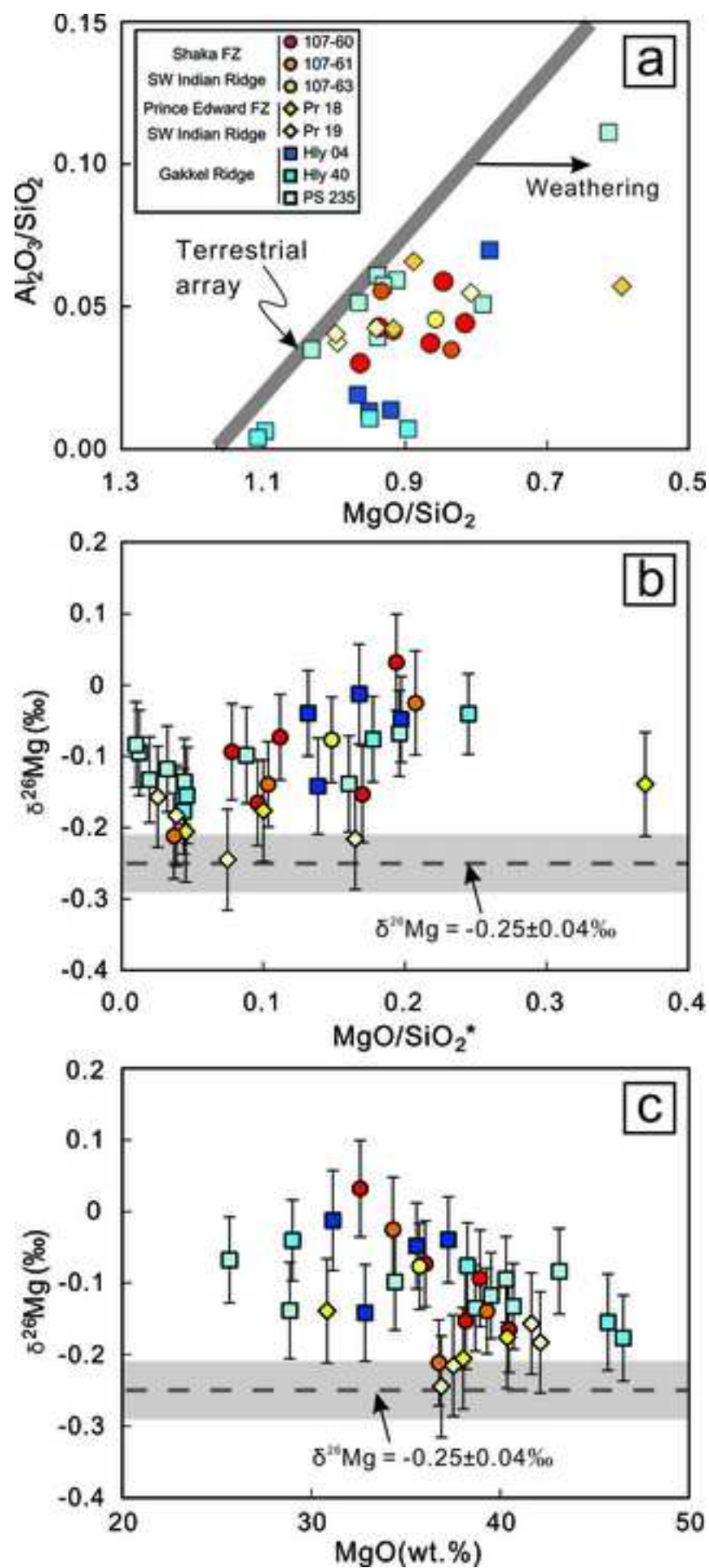


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